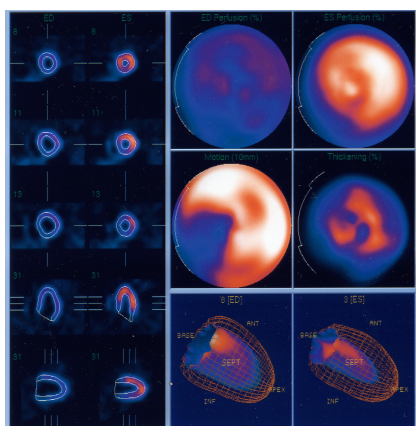


Nuclear Chemistry

Chapter 21

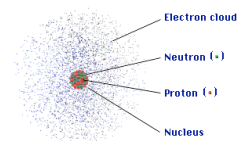
Chemistry 222
Professor Michael Russell

MAR Last update: 4/29/24



MAR

Nuclear Chemistry: the study of the nucleus



Nucleus = neutrons and protons
Differs from "normal" electron chemistry

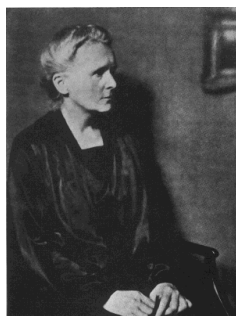
Early Nuclear Chemistry Pioneers:

Marie Curie

Marie Curie
Pierre Curie
Henri Becquerel

Noticed beams of light on photographic plates

Danger (and potential) of nuclear chemistry poorly understood



MAR

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Atomic Number, Z

All atoms of the same element have the same number of protons in the nucleus, **Z**.

13	← atomic number
Al	← symbol
26.9815	← atomic weight

Mass Number, A

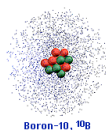
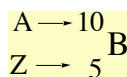
Mass Number

= # protons + # neutrons

A boron atom can have

$$A = 5p + 5n = 10 \text{ and}$$

$$Z = 5 \text{ protons} = 5$$



Isotopes of boron have different # of neutrons but same # of protons (Boron-10, Boron-11)

MAR

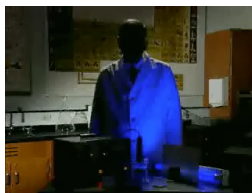
Radioactive Particles

ALPHA	helium nucleus	${}^4_2\text{He}$ or ${}^4_2\alpha$
BETA	electron	${}^0_{-1}\text{e}$ or ${}^0_{-1}\beta$
GAMMA	energy (massless)	γ
PROTON	proton	${}^1_1\text{p}$ or ${}^1_1\text{H}$
NEUTRON	neutron	${}^1_0\text{n}$
POSITRON	antielectron	${}^0_{+1}\text{e}$ or ${}^0_{+1}\beta$

MAR see: *Nuclear Chemistry Guide*

Radioactive Particles

Each particle has different properties



Alpha - stopped by clothes

Beta - stopped by skin

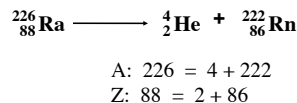
Gamma - stopped by lead

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Balancing Nuclear Reactions

- * Number of **reactant** protons must equal the number of **product** protons
- * Number of **reactant** neutrons must equal the number of **product** neutrons

Example:



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Balancing Nuclear Reactions

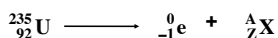
Problem: Uranium-235 decomposes through **beta decay** to a new product. Find the identity of the new product.

Solution:

Uranium has 92 protons, so:



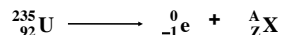
Beta decay means "losing a beta particle", or generating a beta particle as a product, so:



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Balancing Nuclear Reactions

Problem: Uranium-235 decomposes through beta decay to a new product. Find the identity of the new product.

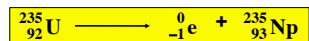


Solution:

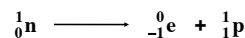
$$235 = 0 + A, \text{ therefore } A = 235$$

$$92 = -1 + Z, \text{ therefore } Z = 93$$

If $Z = 93$, $X = \text{Neptunium (Np)}$, and



The "inner reaction":



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Emission, Decay and Capture

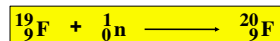
Emission and **Decay** imply a **product** particle

An example of **positron emission**:



Capture implies a **reactant** particle

An example of **neutron capture**:



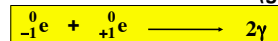
Many types of "particles" and "decays"

MAR

Positrons and Antimatter

Paul Dirac first predicted antimatter in 1928, identical to "regular" matter except for opposite charge

Positrons are **antielectrons**; combining with electrons leads to **annihilation (gamma)**



electron + positron \rightarrow gamma radiation
Used in Positron Emission Tomography (PET)

Many examples of **annihilation reactions** are known:

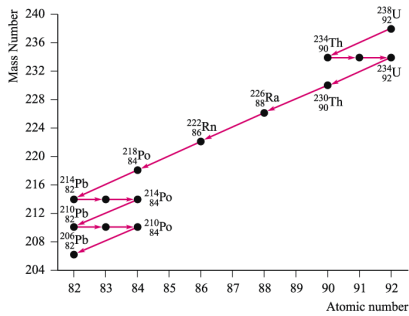


neutrino + antineutrino \rightarrow gamma radiation

Neutrino first postulated by Pauli (CH 221)

MAR

Decay Scheme for Uranium-238



MAR

Metastable Nuclei

When *electrons* are excited in an atom, they relax and emit UV, IR and visible radiation
 A **metastable nuclei** is created when a **nucleus** is excited. Relaxation results in high energy gamma ray emission



metastable Tc \rightarrow "regular" Tc + gamma ray
 Note the "m" for "metastable" and no change in atomic or mass numbers

Metastable nuclei (especially Technetium-99m) used extensively in medical imaging

MAR

The Four Forces of Nature

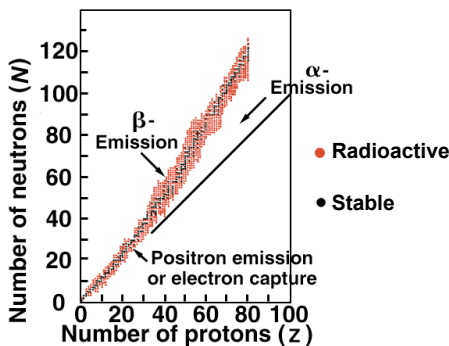
- Gravitation:** weak, long range force
- Weak nuclear:** short range, gives beta decay
- Electromagnetic (E/M):** long range, keeps electrons around nucleus, 10^{36} times more powerful than gravity (aka electrostatic, Coulombic, etc.)
- Strong nuclear:** 10^6 times more powerful than E/M, very short range (10^{-15} m); overpowers E/M repulsion between protons. Strong keeps (protons + neutrons) and (neutrons + neutrons) together. "Glue" that keeps nucleus together

MAR

Nuclear Stability

- Why are nuclei unstable? **E/M Force** (long range) begins overpowering strong force (short range)
- Magnetic dilution** (more neutral neutrons than positive protons) helps stabilize nuclei - to a point
- Up to calcium (Z = 20)**, most stable nuclei occur when **# protons = # neutrons**
- Exceptions:** helium-3 and hydrogen-1
- Up to lead (Z = 82)**, most stable nuclei occur when **# protons < # neutrons**
- Beyond lead, all isotopes unstable and radioactive**

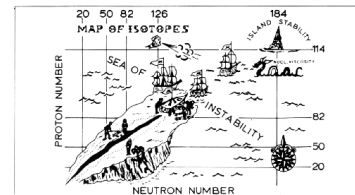
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A "peninsula of stability" (black dots) in a "sea of instability" (red dots)

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An "Island of Stability"?



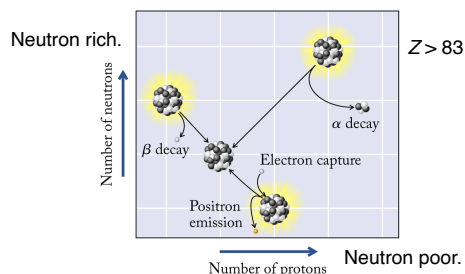
Glenn Seaborg and Ken Moody believe that heavier elements can be made - exciting!

Interested? See: <http://www.pbs.org/wgbh/nova/sciencenow/3313/02.html>

MAR

Nuclear Stability

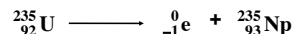
Isotopes *often* decay based on their number of neutrons:



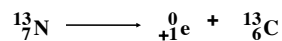
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Nuclear Stability

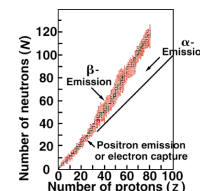
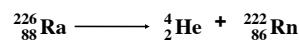
β -decay occurs in isotopes with a high neutron to proton ratio. Result: higher atomic number, same mass number



Low neutron to proton ratio: positron emission or electron capture. Result: lower atomic number, same mass number



All elements beyond Bi decay, usually by α -decay



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Nuclear Stability

Energy required to overcome positive-positive repulsion of protons is substantial - use **strong** (and **weak**) forces

- Strong and weak forces much stronger than electromagnetic force or gravity

Nuclear binding energy, E_b , used to estimate force contribution

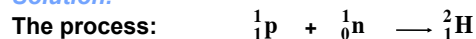
- E_b is the negative of the energy change if nucleus formed directly from individual protons and neutrons

MAR

Nuclear Stability

Problem: Calculate the binding energy, E_b , for deuterium (hydrogen-2).

Solution:



Note that: $\text{mass}_p + \text{mass}_n \neq \text{mass}_D$

$$1.007825 + 1.008665 \neq 2.01410$$

$$2.016475 \neq 2.01410$$

$$\Delta m = -0.00239 \text{ g/mol}$$

mass is *not* conserved!

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Nuclear Stability

Problem: Calculate the binding energy, E_b , for deuterium (hydrogen-2). $\Delta m = -0.00239 \text{ g/mol}$

Solution:

Now use **Einstein's equation:** $\Delta E = \Delta mc^2$

$$\Delta E = (-2.39 \times 10^{-6} \text{ kg mol}^{-1})(2.998 \times 10^8 \text{ m s}^{-1})^2$$

$$\Delta E = -2.15 \times 10^{11} \text{ J mol}^{-1}, \text{ and}$$

$$E_b = -\Delta E, \text{ so}$$

$$E_b = 2.15 \times 10^8 \text{ kJ mol}^{-1}$$

E very negative (exothermic) - lots of energy produced. E_b very positive - lots of energy saved through stabilization

MAR

Nuclear Stability

Problem: Calculate the binding energy, E_b , for deuterium (hydrogen-2). $\Delta m = -0.00239 \text{ g/mol}$

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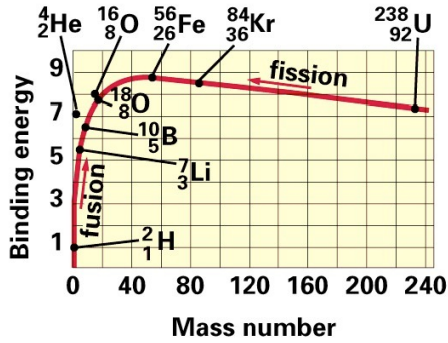
$$E_b \text{ per nucleon} = 2.15 \times 10^8 / (1 \text{ p} + 1 \text{ n}) \text{ nucleons} = 1.08 \times 10^8 \text{ kJ mol}^{-1} \text{ nucleon}^{-1}$$

This is the strong force contribution in the nucleus!

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Nuclear Stability

Can use Binding energy (E_b) to calculate stability of nuclei:



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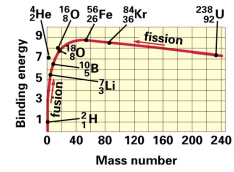
Note that **iron-56** is the most stable element *thermodynamically* - end of solar cycles, etc.

Nuclear Stability

Elements with $Z < 26$ can use **fusion** to become more like iron-56

Elements with $Z > 26$ can use **fission** to become more like iron-56

Kinetics of nuclear reactions important - half life, rate of decay, etc.

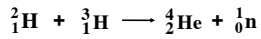


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Tremendous amounts of energy are generated when light nuclei combine to form heavier nuclei - **nuclear fusion**

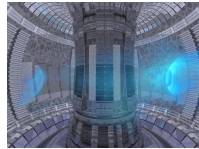
"Clean" energy, used in stars and bombs, requires **plasma** and/or high temperatures - high activation energy barrier

No "meltdown" - reaction just stops, no waste



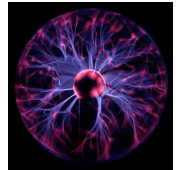
$$E = -1.7 \times 10^9 \text{ kJ/mol}$$

Fusion



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Plasma

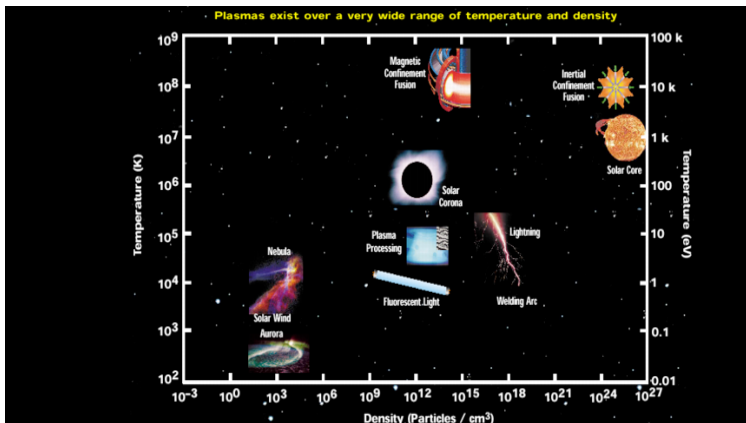


Plasma is the fourth state of matter - along with solids, liquids and gases

Plasma is an electrically conducting fluid composed of freely roaming electrons and ions; strong magnetic and electric fields; hot!

Plasmas comprise the vast majority of the apparent universe, and only in occasional "islands" (like the planet Earth) is matter found in condensed forms (solids, liquids, gases)!

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Plasma - in your microwave?

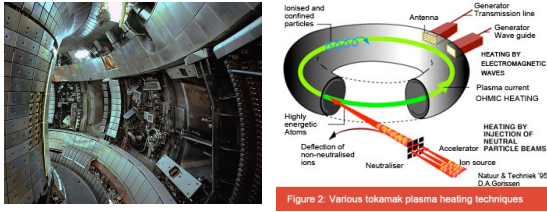


Video by *Sergiu and Ben Todor* 2010

Grape microwave "plasma" - ???

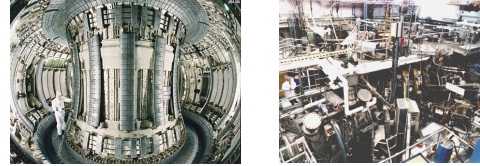
Fusion - Tokamak Reactor

Tokamak reactor uses magnetic fields to constrict plasma for fusion in "donut" shape; most promising "future" magnetic fusion device



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Fusion - Tokamak Reactor

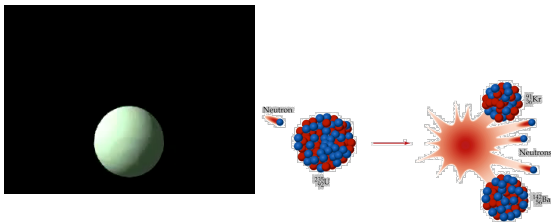


Tokamak reactors - inside and outside views

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Fission

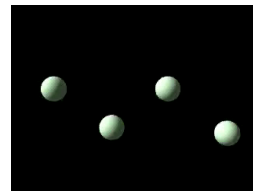
Tremendous amounts of energy (*electricity*) generated when heavy nuclei split to form lighter nuclei - *nuclear fission*
Generally requires a "neutron trigger"



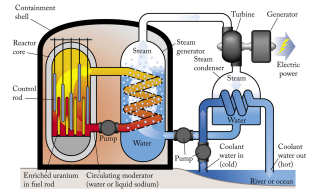
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Excess neutrons *must* be controlled!
Supercritical chain reactions can result without proper caution
Waste products from fission messy and virtually perpetual
More than 400 nuclear fission plants in 30 different countries!

Fission



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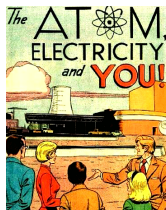


Uses for Nuclear Chemistry

Energy: **Fission** (commercial) and **Fusion** (coming!)

Nuclear fission "problem events":

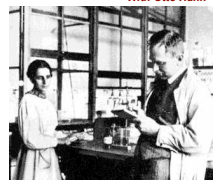
- SL-1, Idaho Falls (1961)
- Three Mile Island, Pennsylvania (1979)
- Chernobyl, Ukraine (1986)
- Fukushima Daiichi, Japan (2011)
- Hanford, Washington (ongoing!)



LISE MEITNER - *unsung hero*



With Otto Hahn

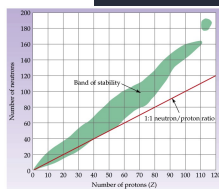
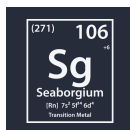
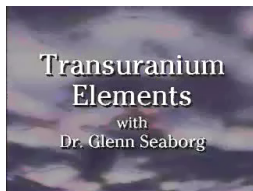


Meitner and colleagues Otto Hahn, Fritz Strassmann & Otto Frisch explained the process of **fission** (1938)
Meitnerium (Mt, #109) named after her
Forced to work in basement, never got Nobel Prize (but Hahn did!)
Pioneering woman in a male-dominated field; deserves more credit for her work

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Uses for Nuclear Chemistry

Expanding the Frontiers of Science



Dr. Glenn Seaborg (1912 - 1999)

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Uses for Nuclear Chemistry

Radiocarbon Dating



$t_{1/2} = 5730$ years for carbon-14
Accurate up to 60,000 years old!

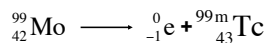
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Uses for Nuclear Chemistry

Medicine: **PET and MRI**



Tc (in PET) not found in nature, created via Mo:



- Tc-99m: metastable; $t_{1/2} = 6.0$ hrs.
- Can be incorporated into variety of compounds to target specific organs (heart, etc.).

PET = Positron Emission Tomography
MRI = Magnetic Resonance Imaging

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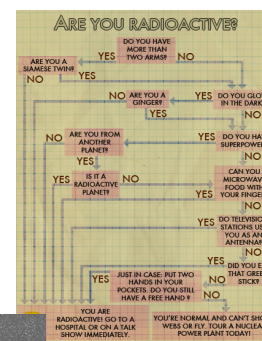
End of Chapter 21

See:

- Chapter Twenty-one Study Guide
- Chapter Twenty-one Concept Guide
- Important Equations (following this slide)
- End of Chapter Problems (following this slide)



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Important Equations, Constants, and Handouts from this Chapter:

- all of the first order kinetics equations apply. See the Nuclear Chemistry Guide (handout)
- decay or emission = product
- capture = reactant
- know how to balance nuclear reactions

$$-E_b = \Delta E = \Delta mc^2$$

c = speed of light = 2.998×10^8 m/s
use kg/mol for Δm

1st Order Integrated Rate Law:

$$\ln \frac{[R]}{[R_0]} = -kt$$

$$t_{1/2} = \frac{0.693}{k}$$

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End of Chapter Problems: Test Yourself

- What particle is emitted when Gold-198 decays to mercury-198?
- What particle is emitted when radon-222 decays to polonium-218?
- What particle is emitted when indium-110 decays to cadmium-110?
- What is emitted when hafnium-178m decays?
- Boron has two stable isotopes, ${}^{10}\text{B}$ and ${}^{11}\text{B}$. Calculate the binding energies per nucleon of these two nuclei. The required masses (in grams per mole) are proton = 1.00783, neutron = 1.00867, boron-10 = 10.01294, and boron-11 = 11.00931.
- Gold-198 is used in the diagnosis of liver problems. The half-life of ${}^{198}\text{Au}$ is 2.69 days. If you begin with 2.8 μg of this gold isotope, what mass remains after 10.8 days?

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End of Chapter Problems: Answers

1. beta particle
2. alpha particle
3. positron particle
4. hafnium-178
5. boron-10: 6.26×10^8 kJ/mol-nucleons; boron-11: 6.70×10^8 kJ/mol-nucleons
6. $0.17 \mu\text{g}$

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